## String Theory I

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## Exercise sheet 1

Due on October 25, 2019

## Exercise 1: Tensors and tensor densities

Consider a coordinate change

$$x^{\mu} \to x^{\prime \mu} \ . \tag{1.1}$$

Recall that tensors can be defined via their transformation properties under (1.1). More concretely, the components of say a (1,1) tensor  $T^{\mu}{}_{\nu}$  transform under (1.1) according to

$$T^{\mu}{}_{\nu} \to T'^{\mu}{}_{\nu} = \frac{\partial x'^{\mu}}{\partial x^{\rho}} \frac{\partial x^{\sigma}}{\partial x'^{\nu}} T^{\rho}{}_{\sigma} . \qquad (1.2)$$

A (1,1) tensor density  $\tilde{T}^{\mu}{}_{\nu}$  of weight w is defined by the transformation property

$$\tilde{T}^{\mu}{}_{\nu} \to \tilde{T}^{\prime \mu}{}_{\nu} = |J|^{w} \frac{\partial x^{\prime \mu}}{\partial x^{\rho}} \frac{\partial x^{\sigma}}{\partial x^{\prime \nu}} \tilde{T}^{\rho}{}_{\sigma} , \qquad (1.3)$$

with  $J = \det\left(\frac{\partial x^{\mu}}{\partial x'^{\nu}}\right)$  (and analogous definitions for tensor densities of different rank).

(a) Given a tensor  $T_{\mu\nu}$ , show that  $(\det T_{\mu\nu})^{1/2}$  is a scalar density of weight 1.

(b) Consider now fields of tensors and tensor densities, e.g.  $T^{\mu}{}_{\nu}(x)$  etc.. An infinitesimal transformation (1.1) can be expanded  $x'^{\mu} = x^{\mu} - \epsilon^{\mu}(x) + \ldots$  Show the following infinitesimal transformations for a scalar field  $\Phi(x)$ , the metric  $g_{\mu\nu}(x)$  and the associated density  $(-\det g_{\mu\nu})^{1/2}$ :

$$\delta \Phi = \epsilon^{\mu} \partial_{\mu} \Phi , \qquad (1.4)$$

$$\delta g_{\mu\nu} = \epsilon^{\gamma} \partial_{\gamma} g_{\mu\nu} + (\partial_{\mu} \epsilon^{\gamma}) g_{\gamma\nu} + (\partial_{\nu} \epsilon^{\gamma}) g_{\gamma\mu} = \nabla_{\mu} \epsilon_{\nu} + \nabla_{\nu} \epsilon_{\mu} (1.5)$$

$$\delta(-\det g_{\mu\nu})^{1/2} = \partial_{\gamma}(\epsilon^{\gamma}(-\det g_{\mu\nu})^{1/2}) . \tag{1.6}$$

The second equality of (1.5) assumes the metric compatible connection, i.e.  $\nabla_{\gamma}g_{\mu\nu} = 0$ . In order to verify that equality, use the Christoffel symbols  $\Gamma^{\mu}_{\nu\rho} = \frac{1}{2}g^{\mu\sigma}(\partial_{\nu}g_{\rho\sigma} + \partial_{\rho}g_{\sigma\nu} - \partial_{\sigma}g_{\nu\rho}).$ 

**Hint:** A scalar field, for instance, is defined via  $\Phi'(x') = \Phi(x)$  and  $\delta\Phi$  in (1.4) is given by  $\delta\Phi(x) \equiv \Phi'(x) - \Phi(x)$ .

## Exercise 2: Equation of motion for a point particle

(a) Consider the variation of the point particle action

$$S = \int_{\tau_i}^{\tau_f} d\tau \mathcal{L}\left(x^{\mu}(\tau), \frac{dx^{\mu}}{d\tau}(\tau)\right) , \qquad (2.1)$$

under a variation  $\delta x^{\mu}(\tau)$  of the particle trajectory. Assume that the variation  $\delta x$  vanishes at  $\tau_i$  and  $\tau_f$ . Show that Hamilton's action principle  $\delta S = 0$  implies the Euler Lagrange equations

$$\frac{d}{d\tau}\frac{\partial\mathcal{L}}{\partial\dot{x}^{\mu}} = \frac{\partial\mathcal{L}}{\partial x^{\mu}} , \qquad (2.2)$$

where  $\dot{x}^{\mu} = \frac{dx^{\mu}}{d\tau}(\tau)$ .

(b) Specify now to the case of a charged particle moving in an electromagnetic field, albeit in flat space. Such a particle is described by the action

$$S = -m \int_{\mathcal{P}} ds + q \int_{\mathcal{P}} d\tau A_{\mu}(x(\tau)) \frac{dx^{\mu}}{d\tau}(\tau) , \qquad (2.3)$$

where

$$ds^{2} = -\eta_{\mu\nu} \, dx^{\mu}(\tau) \, dx^{\nu}(\tau) \; , \qquad (2.4)$$

 $\mathcal{P}$  is the worldline of the particle and the integral along it amounts to an integral from  $\tau_i$  to  $\tau_f$  when the worldline is parameterized by  $\tau$ .

Use (2.2) to show that the equation of motion is

$$\frac{dp_{\mu}}{d\tau} = qF_{\mu\nu}\frac{dx^{\nu}}{d\tau} , \qquad (2.5)$$

where

$$p_{\mu} = m u_{\mu} = m \frac{dx_{\mu}}{ds} , \qquad (2.6)$$

$$F_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu} . \qquad (2.7)$$

(c) Now consider a neutral point particle moving in a curved space-time, i.e.

$$S = -m \int_{\mathcal{P}} ds , \quad ds^2 = -g_{\mu\nu}(x(\tau)) \, dx^{\mu}(\tau) \, dx^{\nu}(\tau) \; . \tag{2.8}$$

Show that the equation of motion is the geodesic equation

$$\frac{d^2 x^{\mu}}{ds^2} + \Gamma^{\mu}_{\nu\rho} \frac{dx^{\nu}}{ds} \frac{dx^{\rho}}{ds} = 0 , \qquad (2.9)$$

where the Christoffel symbols  $\Gamma^{\mu}_{\nu\rho}$  were given at the end of exercise 1.

For questions:

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