

Neutrino Physics Course

Lecture XI

18/5/2021

LMU

Spring 2021



SM and Origin of Mass

(Charged) Fermion Masses

Q. Cartan vs # of invariants

A. Adjoint repn.

$$\Sigma \rightarrow U \Sigma U^+$$

→ diagonal

$$\underline{SU(3)} \quad \Sigma = \begin{pmatrix} a & b \\ b & -(a+b) \end{pmatrix}$$



$$\text{Tr } \Sigma^2, \text{ Tr } \Sigma^3$$

$\text{Tr } \Sigma^4 \neq \text{ independent}$

$$\text{Cartan} = \{T_3, T_8\}$$

$$\hookrightarrow \begin{pmatrix} \text{Tr } T_a = 0 \\ T_a = T_a^t \end{pmatrix}$$

$$\Sigma = c_a T_a \quad \Sigma = \bar{\Sigma}^+$$



$$\text{Tr } \Sigma = 0$$

$$\text{diag} = c_3 T_3 + c_8 T_8$$



SM: starts as ew theory

⇒ theory at origin of mass
(Higgs mechanism)

$$SU(2)_L \times U_Y^{(1)} = G_{SM}$$

$T_a, \quad Y_2$

$A_a \quad B$



$$\bullet i \bar{f} \gamma^\mu D_\mu f = \frac{g}{\sqrt{2}} W_\mu^+ \bar{u}_L \gamma^\mu d_L + h.c.$$

$$f_L = \begin{pmatrix} u \\ d \end{pmatrix}_L$$

$$W_\mu^+ = \frac{(A_1 + i A_2)_\mu}{\sqrt{2}}$$

$$+ e A_\mu \bar{f} \partial^\mu Q_{\text{em}} f +$$

$$+ \frac{q}{c_{\text{em}}} Z_\mu \bar{f} \partial^\mu [T_3 - Q h^\mu \partial_\mu] f$$



$$\bullet \frac{1}{2} (D_\mu \bar{\Phi})^+ (D^\mu \bar{\Phi})$$

$$\bar{\Phi}_{\text{em}} = \begin{pmatrix} 0 \\ v+h \end{pmatrix}$$

$$\bar{\Phi}_{\text{general}} = e^{i G_i / g T_i} \begin{pmatrix} 0 \\ v+h \end{pmatrix}$$

$T_i \begin{pmatrix} 0 \\ v+h \end{pmatrix} \neq 0$

$$T_i = \frac{\sigma_i}{2}$$

"would be" NG bosons

Higgs

$$\bar{\Phi}_{\text{general}} \rightarrow U \bar{\Phi}_{\text{general}} =$$

$$= e^{-i G_i / \hbar T_i} \bar{\Phi}_{\text{general}} = \bar{\Phi}_{\text{res}}$$

$$\boxed{A_{\mu}^{iP} = A_{\mu}^{iH} - \frac{q_{\mu} G^i}{v}}$$

Largitudinal A

$$\Downarrow \quad \bar{\Phi} = \bar{\Phi}_{\text{res}} = \begin{pmatrix} 0 \\ u + h \end{pmatrix}$$

$$= \left(1 + \frac{h}{u}\right) \begin{pmatrix} 0 \\ u \end{pmatrix} = \boxed{\left(1 + \frac{h}{u}\right)} \bar{\Phi}_0$$



$$\frac{1}{2} M_Z^2 \bar{z}_\mu z^\mu \left(1 + \frac{h}{e}\right)^2$$

$$+ M_W^2 \underbrace{W_\mu^+ W^{\mu -}}_{\left(A_1^\mu A_{1\mu} + A_2^\mu A_{2\mu} \right)} \left(1 + \frac{h}{e}\right)^2$$

$$\boxed{M_W = \frac{g}{2} v}$$

$$M_Z = \frac{M_W}{\cos \theta_W}$$

↓

(+ O(h²)

$$M_W^2 W_\mu^+ W^{\mu -} + g M_W h W_\mu^+ W^{\mu -}$$

$$+ \frac{1}{2} M_Z^2 \bar{z}_\mu z^\mu + \frac{g}{\alpha \cos \theta_W} M_Z h \bar{z}_\mu z^\mu$$

Γ $\Gamma_{\text{mass of } Z}$
2 complex

$$h \rightarrow W^+ + (W^- \xrightarrow{*} \text{off-shell})$$

$$m_h = 125 \text{ GeV}$$

$$M_W = 80 \text{ GeV}$$

Counting of states

$$U(1) \quad Q \Phi_0 \neq 0 \Rightarrow H_A \neq 0 \\ (H_A \propto \Phi_0)$$

$$SU(2) \quad T_a \Phi_0 \neq 0 \quad (a=1, 2, 3)$$



$$M_A^a \propto \Phi_0 \quad a=1,2,3$$

$$SU(2) \times U(1) \quad T_a \Phi_0 \neq 0, \quad \nexists \Phi_0 \neq 0$$

$$Q_{\text{em}} \Phi_0 = 0$$



$$Q_{\text{em}} = T_3 + Y_2$$

↓

3 broken generators \Rightarrow

3 massive g.b w^+, w^-, z

1 unbroken \Rightarrow massless A

Theorem:

$$\text{Ta } \Phi_0 \neq 0 \Leftrightarrow \text{Ac}: M_{Aa} \propto \Phi_0$$

$$1 \longleftrightarrow 1$$

gauge symmetry = local
 $\alpha = \alpha(x)$

• global $\alpha = \text{const.}$

• no gauge bosons $D_\mu \rightarrow ?_\mu$

• f_i are real, physical states



$M_{0i} = \odot$ (No medium)

local quantity

$$T_{\alpha} \not{\Phi}_0 \neq 0$$

$$M_{A\alpha} \propto \not{\Phi}_0$$

global quantity

$$\overline{T}_{\alpha} \not{\Phi}_0 \neq 0$$

$$M_{G\alpha} = 0$$

① Prove the eq

$$M_A \neq 0, (+ \underline{\theta})$$

$$g A_\mu \bar{\psi}_L \gamma^\mu \psi_L \quad (\underline{w} \not{\gamma}_R)$$

$(g \rightarrow 0) \Rightarrow$ decoupling !

① Higgs theory

$$\mathcal{L} = \frac{1}{2} (\partial_\mu \phi)^* (\partial^\mu \phi) - V(\phi)$$

$$- \frac{1}{4} F_{\mu\nu} F^{\mu\nu} - \left(g \bar{\psi}_L \phi \psi_R + h.c. \right)$$

$$+ i \bar{\psi}_L \gamma^\mu D_\mu \psi_L + i \bar{\psi}_R \gamma^\mu D_\mu \psi_R$$

$$D_\mu \psi_L = (\partial_\mu - ig A_\mu) \psi_L \quad D_\mu \psi_R = \partial_\mu \psi_R$$

$$D_\mu \phi = (\partial_\mu - ig A_\mu) \phi$$

$$\phi = v + h + iG$$

↓
scattering

$$\alpha g^2 \left(\begin{array}{c} \text{wavy line} \\ A \end{array} \right) + \left(\begin{array}{c} \text{---} \\ G \end{array} \right) \propto y^2$$

(a)

(b)

$$(\zeta=1) \quad \downarrow \quad \Downarrow \quad \quad \quad = \text{independent of } \zeta$$

$$\frac{g (\bar{\psi}_L \gamma^\mu \psi_L)^2}{q^2 - M_A^2} + g^2 \frac{m_f^2 / m_A^2}{q^2 - M_A^2} (\bar{\psi}_R \psi_R)^2$$

($\zeta = 1$ gauge)

$$D(6) \propto \frac{1}{h^2 - M_A^2}; \Delta_{\mu\nu} \propto \frac{g_{\mu\nu}}{h^2 - M_A^2}$$

$$D(G) = \frac{i}{k^2 - g M_A^2} \quad \uparrow$$

$$\Delta_{\mu\nu} = \frac{-i}{k^2 - M_A^2} \left[g_{\mu\nu} + (\beta - 1) \frac{k_\mu k_\nu}{k^2 - g M_A^2} \right]$$

$$(b) \quad y \left[\bar{\psi}_L (\nu + h + iG) \psi_R + \bar{\psi}_R (\nu + h - iG) \psi_L \right]$$

$$= y(\nu + h) (\bar{\psi}_L \psi_R + \bar{\psi}_R \psi_L) + iGy (\bar{\psi}_L \psi_R - \bar{\psi}_R \psi_L)$$

$$= y(\nu + h) \bar{\psi} \gamma - iGy \bar{\psi} \gamma_5 \gamma$$



$$m_f = y\nu, \quad M_A = g\nu$$

$$y = g \frac{u_f}{M_A}$$

$$\Rightarrow h \bar{f} f g \frac{u_f}{M_A}$$

local \longrightarrow global

$$M_A \neq 0 \xrightarrow{(?)} u_Q = 0$$

$$D_\mu = \partial_\mu - i g Q A_\mu$$



$$f = 0 \text{ limit}$$

$$g = 0$$

$|a| \rightarrow 0$, $|b| \neq 0$

$$A \rightarrow A_b \xrightarrow{\cancel{\frac{g^2 m_f^2}{g^2 v^2}}} \frac{1}{q^2} (\bar{\psi} \gamma_5 \psi)^2$$

↑
 massless particle =
 $= N \& boson$

(i) \exists massless G ($m_a=0$)

(ii) has $\bar{\psi} \gamma_5 \psi$ int. !!!

$$(i_{ii}) \quad y = \frac{m_f}{v}$$

A decouples when $v \rightarrow \infty$

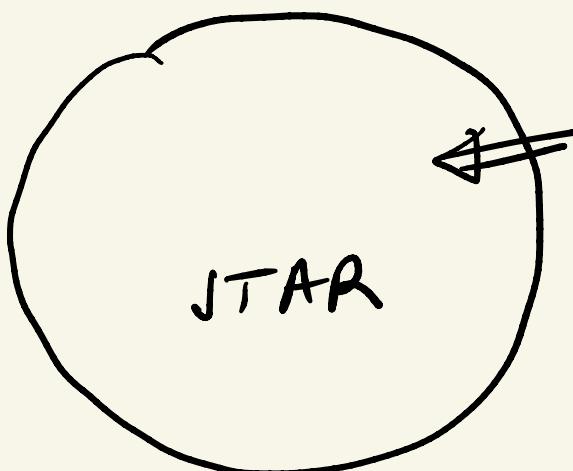
v = "cut-off"

massless message G



charge gravity! ?

($V_{qr}(r) \Leftarrow$ charged?)



many protons +
neutrons

$$N_0 \approx 10^{57}$$

= heavy ($M = N \text{ GeV}$)

• $\bar{\gamma} \gamma_5 \gamma G$

\not{t} spin-dependent

NR limit

$$\gamma = \begin{pmatrix} u \\ \frac{\vec{p} \cdot \vec{\sigma}}{2m} u \end{pmatrix} \text{ small}$$

$$\left. \begin{array}{l} \gamma^0 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \\ \gamma_5 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \end{array} \right\}$$

$$\gamma^\mu p_\mu \gamma = m \gamma$$

$$= \begin{pmatrix} E-m & \vec{p} \cdot \vec{\sigma} \\ -\vec{p} \cdot \vec{\sigma} & E+m \end{pmatrix} \begin{pmatrix} u \\ v \end{pmatrix} = 0$$

$$v = \frac{\vec{p} \cdot \vec{\sigma}}{E+m} u \approx \frac{\vec{p} \cdot \vec{\sigma}}{2m} u$$

$$\text{gyr} \propto u^+ \frac{\vec{\sigma} \cdot \vec{p}}{2m} u =$$

$$= \underbrace{\left(u^+ \frac{\vec{\sigma}}{2} u \right)}_{\langle \vec{s} \rangle} \cdot \vec{p} = \langle \vec{s} \rangle \cdot \vec{p}$$

$$\langle \vec{s} \rangle_{\text{nuclei}} \approx O(1)$$

$$\langle \vec{s} \rangle_{\text{star}} \approx O(1)$$

spin is not coherent!

G exchange = negligible

compared to gravity

↓

Sun "radiates" δ

$$\Rightarrow \boxed{v \gtrsim 10^{10} \text{ GeV}}$$

↑

Red giants

↓

$\boxed{\text{back to SM}}$

$$q_L \equiv \begin{pmatrix} u \\ d \end{pmatrix}_L \quad u_R \quad d_R$$

$$l_L \equiv \begin{pmatrix} \nu \\ e \end{pmatrix}_L \quad e_R$$

$$d\bar{y} = \bar{q}_L^T y_d \bar{\Phi} d_R + \bar{l}_L^T \bar{\Phi} y_e e_R$$

$$+ \bar{q}_L q_R i \gamma_5 \vec{\Phi}^* u_R + h.c.$$

$$\vec{\Phi}_u = \begin{pmatrix} 0 \\ v+h \end{pmatrix}$$

$$\Rightarrow \mathcal{L}_Y = \left[\bar{d}_L q_d d_R + \bar{e}_L q_e e_R + \right. \\ \left. + \bar{u}_L q_u u_R + h.c. \right] (v+h)$$

$$= g_v \bar{f} f + g_h \bar{f} f$$

↔

$$m_f = v y \quad y = \frac{m_f}{v} = \frac{g}{2} \frac{m_f}{M_W}$$

$$(M_W = \frac{g}{2} v)$$



$h \rightarrow f\bar{f}$ probe directly

$$\Gamma(h \rightarrow f\bar{f}) = \frac{g^2}{8\pi} m_h \quad (m_h \gg m_f)$$

$$\propto \left(\frac{m_f}{M_W}\right)^2 m_h$$

Origin of mass:

probing $h \rightarrow f\bar{f}$

LHC

w^+, w^-, z }
 t, b, τ } Higgs

• e, u, d ($m \rightarrow 0$)

{ have to probe Higgs!

• μ, c, s



Direct test of Higgs

Universe was hot

$$T \gg v \simeq M_W \simeq 100 \text{ GeV}$$

$$(M_W = \frac{g}{2} v)$$

$$g \simeq 0.6$$

$$e = g \hbar u \theta_w$$

$$e = \frac{1}{3}^2 \quad \theta_w \approx 30^\circ$$

Weinberg '74

$$V_T = V_{T=0} + a T^2 \Phi^+ \bar{\Phi}$$

$$a = g^2 + \lambda + |g|^2 > 0$$

$$\text{high } T \Rightarrow \Phi_0(T) = 0$$

$$V_{T=0} = \frac{\lambda}{9} (\Phi^+ \bar{\Phi} - v^2)^2$$

$$= \frac{\lambda}{9} (\Phi^+ \bar{\Phi})^2 - \frac{\lambda v^2}{2} \Phi^+ \bar{\Phi} + \text{const.}$$

sym. breaking

$$a T^2 \gg -v^2 \lambda \Rightarrow \langle \Phi_0 \rangle_T = 0$$

$$T \simeq 100 \text{ GeV} \simeq 10^{13} \text{ eV} = 10^5 \text{ MeV}$$

early universe

Origin of mass (f)

$$h \bar{f} f \frac{m_f}{v} \not\propto$$

$$\Rightarrow \Gamma(h \rightarrow f \bar{f}) \propto m_f^2$$

NOT asking: why $m_f \rightarrow$
 m_f ?